

Resonant Origins for Pluto's High Inclination

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Goals

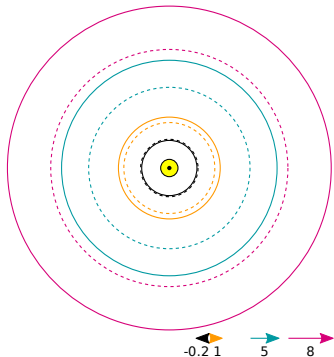
Explain Pluto's high eccentricity ($e = 0.24$) and high inclination ($i = 17^\circ$) using resonances

- Three candidates
 - 1 6:4 mean motion resonance
 - 2 1:1 secular resonance
 - 3 2:1 secular resonance

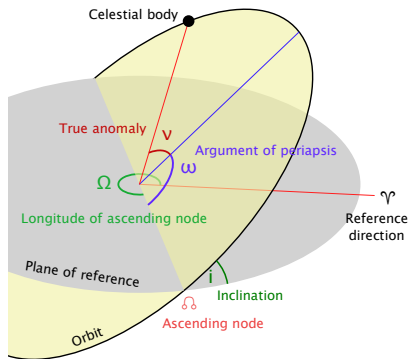


Planetary Migration by Scattering Planetesimals

- Planets other than Jupiter preferentially scattered planetesimals inward, migrated outward
- Migrations move locations of resonances, catching Pluto
- If migration rate is slow enough, characteristic effect on resonances is rate-independent



Orbital Elements & Symmetries



Orbital Elements: a , e , i , Ω ,
 $\varpi = \Omega + \omega$, λ ($\dot{\lambda} \approx n$)

Secular Variables

$$h = e \sin(\varpi)$$

$$k = e \cos(\varpi)$$

$$p = \sin(i/2) \sin(\Omega)$$

$$q = \sin(i/2) \cos(\Omega)$$

Eigenfrequencies: f , g



Resonant Behavior

Mean Motion Resonance

Simple ratio of orbital periods
(dependent on λ , n)

Secular Resonance

Simple ratio of precession
periods (averaged orbits)

- Form resonant arguments subject to symmetries
 - Good: $6\lambda_P - 4\lambda_N - 2\Omega_P$, $2\Omega_P - \Omega_N - \Omega_J$
 - Bad: $3\lambda_P - 2\lambda_N - \Omega_N$, $2\Omega_P - \Omega_N$
- Capture
- Jump



Simulation and Analysis

Both of pre-existing and new software used throughout project.

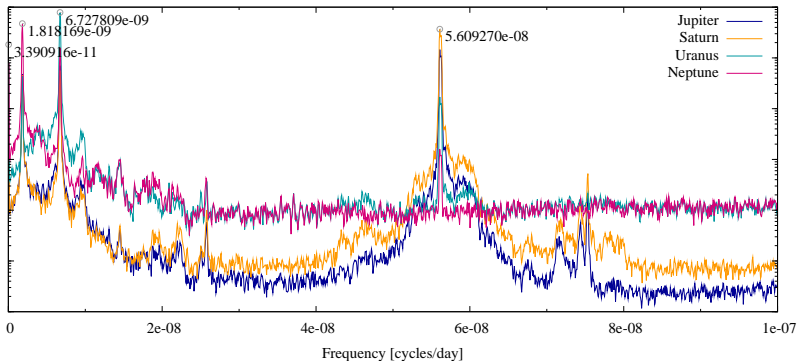
- Used *HNBody* and *HNDrag* to simulate Solar System over billions of years (> 24 GB of data generated)
- To determine secular eigenfrequencies, wrote code to perform FFT on orbital elements
- Features of *PowerSpectrumEstimator*:
 - Data windowing to reduce spectral leakage
 - Overlapping data segments to minimize variance
 - Automatic peak finding with inverse quadratic interpolation
 - Removal of aliased peaks
 - Orthogonality of total angular momentum



Example: Outer Solar System p Spectra

Matches g_6 to better than 1%; matches g_7 to within 7%; matches g_8 to within 25%; g_5 is effectively 0

Spectra of 'p' for the Outer Solar System



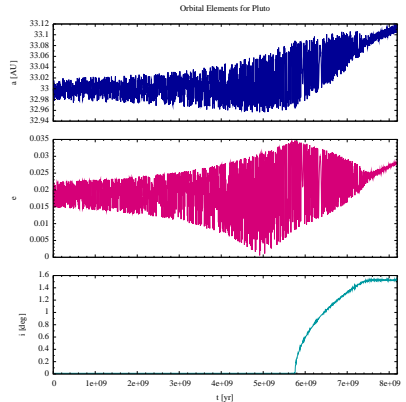
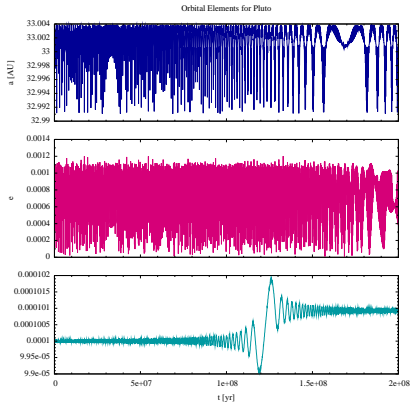
Candidate #1 – 6:4 Mean Motion Resonance

Pluto is currently trapped in a 3:2 eccentricity resonance ($3n_P - 2n_N - \dot{\omega}_P$) and a Kozai resonance ($\dot{\Omega}_P - \dot{\omega}_P$). Together, these imply a 6:4 inclination resonance ($6n_P - 2n_N - 2\dot{\Omega}_P$).

- Initially, these were split (no Kozai resonance)
- Being first-order, eccentricity resonance is stronger
- Simulations rule out capturing in inclination resonance first
- What about afterwards?



Examples of Mean Motion Resonances



Migration rates too slow, inclination rise too small



Candidate #2 – 1:1 Secular Resonance

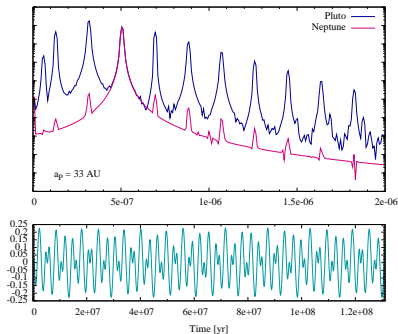
A 1:1 resonance ($\dot{\Omega}_P - \dot{\Omega}_N$) should be easier to find and more powerful than a 2:1 resonance.

- Studied an idealized Jupiter+Neptune+Pluto system
- May have been present at Solar System formation
- Could capture into 3:2 mean motion resonance at just the right time, maintain high inclination

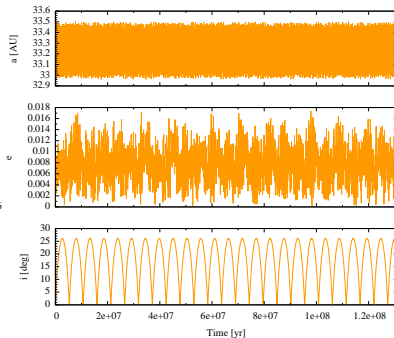


Example of 1:1 Secular Resonance

Plot and Spectrum of Pluto's 'p'



Orbital Elements for Pluto



Static inclination resonance extremely broad and powerful
(3 AU, 25°)

Secular Resonances in the Solar System

In full Solar System, 1:1 resonance is not as broad or powerful. Still, migrating across makes jump or capture possible.

- Inclination jump of 10° observed near initial conditions
- Capture raises more questions: when/how did it break out?
- Leaves observed 2:1 resonance a coincidence

Early proximity to 1:1 indicates that 2:1 was not active prior to capture in eccentricity resonance.



Candidate #3 – 2:1 Secular Resonance

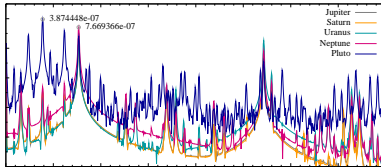
By raising $M_U \rightarrow 1.8M_U$, we could create conditions where $2p_1 \approx g_8$. By dragging Pluto directly, we could study strength of jump and capture.

- Raising $M_U \iff$ increasing Uranus's initial position
- Jump is too weak (2°) to explain current inclination
- What about capture?

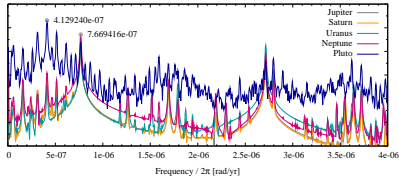


Example of 2:1 Capture

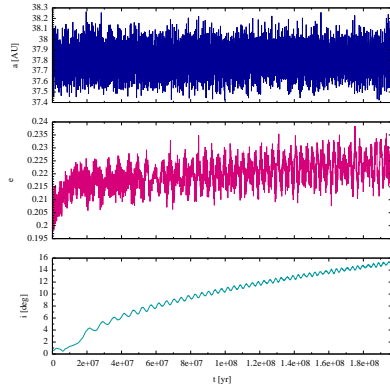
Spectra of γ' for the Outer Solar System: Initial



Spectra of γ' for the Outer Solar System: Resonance (2e7 yr)



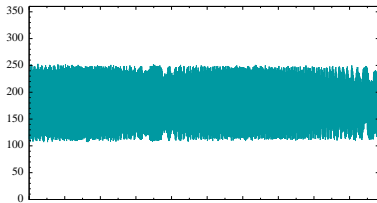
Orbital Elements for Pluto



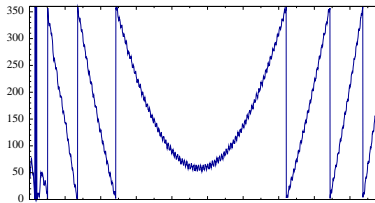
Capture is possible! Yields $i \rightarrow 16^\circ+$

Active Resonances in 2:1 Capture

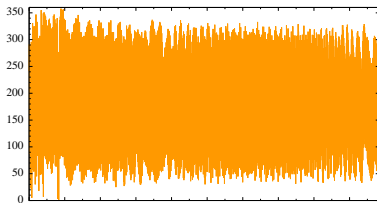
$$3\lambda_p - 2\lambda_N - \varpi_p$$



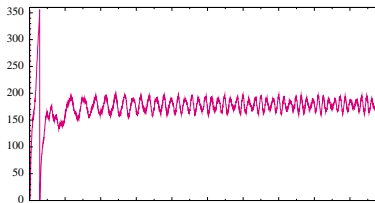
$$2\Omega_p - \Omega_N - 0$$



$$6\lambda_p - 4\lambda_N - 2\Omega_p$$



$$2\Omega_p - 2\varpi_p$$



Summary

Currently, no overwhelmingly likely explanation. However, some can be ruled out while others can be constrained.

Resonance	Grade	Pros	Cons
Mean Motion	D	Currently active	Could not capture Too weak
Secular 1:1	B	Strong enough Possibly active in early solar system	Not active today Large jump instead of capture
Secular 2:1	B+	Possibly active today Capable of capture	$M_U \rightarrow 1.8M_U$ Dragging Pluto, not Neptune

